



# Holography and hydrodynamics

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**Abstract** We describe some recent insights from string theory into holography, especially in the context of a hydrodynamic description of black hole physics

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## 1. Introduction

Gravity seems to stand apart from all the other forces of nature in many different ways. Whereas the quantum mechanics of the other forces are beautifully captured within the framework of local quantum field theory, gravity has resisted all such attempts to tame it. Partly, the reason lies from Einstein's observation that gravity is woven into the very fabric of spacetime. This suggests that to understand the dynamics of gravity might be radically different from that of other forces.

The idea that gravity may be *holographic* is one interesting way in which this difference may manifest itself. This notion was proposed by 't Hooft [1] and elaborated by Susskind [2] in the early nineties to make sense of some aspects of black holes. It remained a tantalising idea until a very concrete realisation was proposed by Maldacena, for a theory of gravity in an asymptotically Anti-de Sitter (AdS) spacetime. The ramifications of the Maldacena conjecture for the holographic nature of gravity continue to be unearthed and we seem to be uncovering a very beautiful picture.

In this article I will try and briefly sketch a few of the relevant notions that have emerged in the last few years. I will focus on some aspects of black holes that are holographically reflected in the dynamics of strongly interacting thermal field theories. In particular, we will

describe how the *hydrodynamic* behaviour of such gauge theories is encoded in properties of black hole horizons. Not only has this led to insights into the physics of field theories but it is also giving a new perspective on the membrane paradigm for black holes.

AKR, I believe, would have been very happy to see some of these developments since they provide fresh illumination on some of the concepts that AKR himself helped develop. This article is dedicated to his memory.

## 2. Holography

The universal Bekenstein-Hawking expression for the entropy of black holes

$$S_{BH} = \frac{A_H}{4G_N} \quad (1)$$

was one of the first clues to the holographic nature of gravity. Recall that  $A_H$  is the *area* of the horizon and  $G_N$  is Newton's constant. The area dependence indicates that any microscopic accounting of this entropy in terms of the number of degrees of freedom of some underlying theory must have an unusual property. It cannot be extensive in the volume. In fact, the area dependence seems to say that the degrees of freedom must be more like those of a local field theory in one lower dimension. The bold proposal of 't Hooft was that the *entire gravitational dynamics* in  $d$

*spacetime dimensions is completely encoded in a local theory in  $(d - 1)$  dimensions*. Thus the origin of the term "holography"

The 1997 AdS/CFT conjecture of Maldacena [3] offered the opportunity to flesh out this proposal. The conjecture arose from developments in string theory and describes an exact duality between closed superstring theory in the ten dimensional  $AdS_5 \times S^5$  spacetime and the maximally supersymmetric Yang-Mills theory in four dimensions. The conjecture has since been generalised and is now understood to be valid for a large class of asymptotically  $AdS_5$  spacetimes. In all cases, a semiclassical gravitational description of the spacetime dynamics is captured by the strong interaction dynamics of a 4d field theory.

The connection between dynamics in five dimensional AdS spacetime and four dimensional field theory is a reflection of the holographic nature of this duality. In fact, the 4d spacetime in which the field theory lives is the asymptotic boundary of the  $AdS_5$  spacetime. There is then a precise dictionary [4,5] between observables in the gauge theory on the *boundary* and those of the gravity theory in the *bulk*. In the semiclassical limit for gravity, this dictionary takes the form

$$\int [D\phi] e^{-S_{\text{grav}}[\phi]} = \int [D\phi] e^{-S_{\text{gauge}}[A]} \int_{d^4x_0} \phi(x_0) \mathcal{O} \quad (2)$$

Here the left hand side is the gravity side in which we perform an integral over bulk fields  $\phi$  (including the metric) taking specified values  $\phi_0$  on the boundary. The right hand side is the gauge theory in which we have the path integral over gauge fields  $A$ . The boundary value  $\phi_0$  acts as a source term for a gauge invariant operator  $\mathcal{O}_\phi$  of the field theory. Thus the gravity theory may be viewed as the generating functional for correlators in the boundary gauge theory.

What does this holographic description tell us about black holes, the original motivation of 'tHooft and Susskind? One finds that the Schwarzschild black hole in  $AdS_5$ , for instance, is described by the same gauge theory but now at a finite temperature, which is the Hawking temperature of the black hole [6]. Thus we can translate questions about black hole physics (in five dimensions) into questions in the strongly interacting dynamics of hot gauge theories (in four dimensions) and *vice versa*!

For instance, the equilibrium thermodynamics of the gauge theory describes black hole thermodynamics. In particular, the Bekenstein-Hawking entropy of black holes

is the same as the entropy of the thermal gauge theory in the deconfined phase. The latter, being extensive in the three spatial dimensions is actually proportional to the area of the horizon of the five dimensional black hole (which is a three dimensional sphere). Other thermodynamic features such as phase transitions in the gauge theory, also have their reflection in the gravity side.

This connection to strongly interacting gauge theory took on a new significance because of recent experimental results from the Relativistic Heavy Ion Collider (RHIC) at Brookhaven. In these experiments, the high energy collisions of gold nuclei seem to have created a new temperature phase of deconfined strongly interacting quarks and gluons (see [7,8] for overviews). This strongly interacting Quark-Gluon Plasma (SQGP) is very difficult to describe using the tools of perturbative gauge theory. It appears as if the dual description in terms of gravitational physics of five dimensional AdS black holes is better at reproducing some of the features that the RHIC data exhibit!

### 3. Hydrodynamics

One of the challenges of understanding the RHIC data was in describing phenomena which go somewhat beyond equilibrium dynamics. In fact, the process of formation and evolution of the SQGP phase are best described by hydrodynamics. Remarkably, it appears that the SQGP behaves almost like a perfect fluid. To a good approximation, it is as if the shear viscosity  $\eta$  is negligible. The dimensionless parameter

$$\frac{\eta}{s} = \frac{1}{10}$$

where  $s$  is the entropy per unit volume. If this description of the data is indeed correct then this would be the most perfect fluid observed in nature.

Such a small value for the shear viscosity is difficult to explain using perturbative calculations in the gauge theory. These give a behaviour

$$\frac{\eta}{s} \sim g_{YM}^4 \ln \frac{1}{g_{YM}^2}$$

for small values of the gauge coupling  $g_{YM}$ . In the perturbative calculations, the shear viscosity is extracted from the imaginary frequency pole in the two point function of components of the energy momentum tensor.

Using the *AdS/CFT* dictionary for gauge theory correlators mentioned in the previous section, this two point function has a gravity interpretation in the five dimensional black hole geometry. It can be related to the absorption cross section for polarised gravitons by the black hole [9]. This latter calculation yields a simple prediction for the shear viscosity at very strong coupling

$$\frac{\eta}{s} = \frac{1}{4\pi} \quad (5)$$

Remarkably, this is very close to the value that best fits data.

In fact, there seems to be something universal about this value of the viscosity extracted from gravity calculations. For a large class of black holes, the absorption calculations yield this same value, somewhat like the universality of the area formula for the entropy. It has been conjectured that the above value for  $\eta/s$  is a universal lower bound. See [10] for an overview and references to the literature on these topics.

Viscosity is a familiar notion in a field theory, but does it have a direct interpretation in the gravity side? In fact, physicists studying black holes in the 1970's came up with the notion of a membrane paradigm. The horizon of the black hole, as viewed by an outside observer, appears to have all the properties of a physical membrane. Among other things the membrane has an energy momentum tensor which is that of a fluid. Essentially, this follows from the behaviour of geodesics in the vicinity of the horizon. The shear, rotation *etc.* of these geodesics, as introduced by AKR, are instrumental in the identification with a fluid. The respective coefficients then give the various viscosity coefficients. Amazingly, the ordinary 4d Schwarzschild black hole also has a ratio of  $\eta/s = 1/4\pi$  in this membrane paradigm.

It therefore seems that the hydrodynamics of the gauge theory is holographically capturing the membrane description of black hole horizons. We have grown familiar with the notion of associating an entropy to black hole horizons and now understand it as the entropy of the microscopic gauge description of the system. Similarly, we are now viewing the hydrodynamic properties that can be attributed to the horizon as those of the hydrodynamics of the gauge theory. Thus the holographic gauge theory description gives a physical meaning and sharper picture to the ideas of the membrane paradigm.

There is an aspect of this connection that the attentive reader might have found puzzling. We had mentioned

earlier that the gauge theory is to be viewed as living on the asymptotic boundary of the AdS spacetime. Whereas we are now connecting its properties with those of the horizon which is a different four dimensional surface in the interior of the spacetime. While we don't understand this sharply enough, it appears that long wavelength phenomena of the gauge theory (such as hydrodynamics) capture the dynamics near the horizon. The boundary description is more appropriate as a UV or microscopic one. This dichotomy is a reflection of the fact that the radial direction in the AdS spacetime plays the role of the energy scale in the gauge theory, with asymptotic infinity being the UV and the deep interior near the horizon being the IR.

To summarise, the realisation of holography in AdS spacetimes has enabled us to get a better handle on many aspects of this miraculous connection. The understanding of hydrodynamic phenomena is a step beyond the conventional black hole thermodynamics and shows us how far reaching the connection between the gauge and gravity theories are. It has already provided qualitative and probably quantitative understanding of the SQGP phase discovered at RHIC. The comprehensive nature of the connection promises to also shed light on far from equilibrium situations such as in black hole formation. This has yet to be explored.

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